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Symmetry, Selection Rules, and Energy Levels of
 $\text{Pr}^{3+}\text{YAl}_5\text{O}_{12}$

by John B. Gruber
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Gregory A. Turner

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1. Introduction

The optical properties of trivalent rare-earth ions (R^{3+}) in yttrium aluminum garnet, $Y_3Al_5O_{12}$ (YAG), continue to be of great interest following recent developments in solid-state lasers and magneto-optical materials [1–5].* Over the years, the optical spectrum of Pr^{3+} :YAG has been studied by various groups [6–11], but there remain many uncertainties and ambiguities regarding the assignments. Although Pr^{3+} ions substitute predominantly for Y^{3+} ions in the dodecahedral sites of D_2 symmetry, several minority sites are occupied as well. This observation is not unique to Pr^{3+} :YAG, but is found for other R^{3+} :YAG materials as well [12–15]. Different sites can arise from defects caused by loss of oxygen from the lattice during crystal growth [16]. Consequently, increased attention has been placed on growing high-quality R^{3+} :YAG crystals.

We have investigated the spectroscopy of Pr^{3+} :YAG between 2,000 and 25,000 cm^{-1} at several temperatures between 1.6 and 90 K. In addition to absorption spectra, fluorescence from 3P_0 and site-selective emission spectra from 1D_2 were recorded. The most intense spectra, representing Pr^{3+} ions in D_2 sites, were analyzed according to electric-dipole selection rules for D_2 symmetry. The intense spectra included a number of hot bands that were analyzed. The analyses were aided by the knowledge that the ground state has Γ_3 symmetry and that the 3P_0 level has Γ_1 symmetry [8]. Fifty-one experimental Stark levels whose symmetry labels were established by this method of analysis were compared with a calculation that involved a Hamiltonian containing Coulombic, spin-orbit, and crystal-field terms in D_2 symmetry for all manifolds of the $4f^2$ configuration. The initial set of nine crystal-field parameters, B_{km} , was established from a lattice-sum calculation. By varying the crystal-field parameters, it was possible to obtain overall agreement between calculated and experimental levels with an rms deviation of 11 cm^{-1} . Additional levels, whose symmetry could not be established from experiment, were assigned from the results of the crystal-field calculation.

*References are listed at the end of the report.

2. Experimental Introduction

Two single crystals of $\text{Pr}^{3+}:\text{YAG}$ were grown parallel to the $\langle 111 \rangle$ direction using the Czochralski method. The garnet melt was doped with praseodymium oxide, and the crystal growth took place in a nitrogen atmosphere containing 1000 ppm of oxygen. Discs were cut parallel to the (111) plane. The doping appeared to be uniform throughout the discs. Based on the distribution coefficient for the dopant, the crystals contained approximately 0.02 and 0.08 at. wt.% praseodymium based on yttrium. The concentrations as determined by electron-beam microprobe analysis were 0.018 and 0.072 at. wt.%, respectively.

A Nicolet model 7199 Fourier transform infrared (FTIR) spectrometer was used to obtain spectra between 1500 and 6000 cm^{-1} . The accuracy in determining the wavelength of the absorption peaks was limited by a combination of instrument resolution (0.25 cm^{-1}) and observed spectral linewidths. Absorption spectra between 2.5 and $0.4\text{ }\mu\text{m}$ ($4,000$ to $25,000\text{ cm}^{-1}$) were measured with a Cary model 17D spectrophotometer. At $0.4\text{ }\mu\text{m}$, the accuracy is better than 0.2 nm ; the resolution is better than 0.1 nm . The precision in measuring the separation between peaks having apparent spectral linewidths of 0.2 nm is better than 0.02 nm . For data recorded at the same wavelength on both instruments, agreement is better than 2 cm^{-1} .

A conduction dewar filled with liquid nitrogen or liquid helium was used to obtain crystal spectra at nominally liquid nitrogen or liquid helium temperatures. At least 30 minutes was allowed for equilibration before spectra were obtained. Sample temperatures were not measured directly. However, from previous experience in taking similar spectra, we estimate the crystal temperatures as 90 and 15 K, respectively.

Absorption and emission spectra were also recorded between 0.4 and $0.8\text{ }\mu\text{m}$ using a 1.0-m Czerny-Turner Jarrell-Ash double monochromator with a resolution of 0.02 nm . The measured sample temperature in a liquid helium immersion cryostat was 1.6 K . Wavelengths were calibrated against the spectrum of neon. The wavelengths of the sharpest absorption peaks were determined with an accuracy of 0.05 nm .

3. Symmetry, Selection Rules, and Assignments

Yttrium aluminum garnet has cubic space group symmetry O_h^{10} ($1a3d$) with eight formula units per unit cell [17,18]. The site symmetries of the ions are as follows: D_2 (Y^{3+}), C_{3i} (Al_1^{3+}), S_4 (Al_2^{3+}), and C_i (O^{2-}). The Pr^{3+} ions that substitute for Y^{3+} ions on the dodecahedral sites experience a crystal field of D_2 symmetry. There are six magnetically inequivalent orientations of these sites in the crystal. The local x , y , and z axes of the Pr^{3+} ion sites are oriented along the (110) , $(1\bar{1}0)$, and (001) crystal axes (and the six equivalent directions) with a z -axis coinciding with the (001) crystal axis by convention [19].

The D_2 point group contains four one-dimensional irreducible representations, Γ_1 , Γ_2 , Γ_3 , and Γ_4 according to Bethe notation [20]. For Pr^{3+} ($4f^2$), each $^{2S+1}L_J$ manifold is split into $2J + 1$ nondegenerate components (Stark levels). Table 1 gives the representations for each value of J up to 6. Table 2 presents the selection rules for electric-dipole transitions.

The dipole nature of the electronic transition is difficult to establish in crystals of R^{3+} :YAG because the overall symmetry is cubic. From linear dichroism measurements on samples with unequal site occupation, van der Ziel et al [8] observed only electric-dipole transitions in Pr^{3+} :YAG. They determined that the ground state has Γ_3 symmetry and the second excited Stark level at 50 cm^{-1} has either Γ_2 or Γ_4 .

Table 1. Full rotation compatibility table for D_2 group

J	Γ_1	Γ_2	Γ_3	Γ_4
0	1	0	0	0
1	0	1	1	1
2	2	1	1	1
3	1	2	2	2
4	3	2	2	2
5	2	3	3	3
6	4	3	3	3
Total ^a	28	21	21	21

^aTotal number of Γ_n in electronic configuration $4f^2$.

Table 2. Electric-dipole selection rules for D_2 symmetry

	Γ_1	Γ_2	Γ_3	Γ_4
Γ_1	—	y	z	x
Γ_2	y	—	x	z
Γ_3	z	x	—	y
Γ_4	x	z	y	—

symmetry. We chose Γ_4 symmetry for the level at 50 cm^{-1} on the basis of our crystal-field splitting calculation and the consistency we found in our analysis of hot-band data, which included transitions from the 50-cm^{-1} level.

For D_2 symmetry, electric-dipole transitions from the ground state to excited Stark levels having Γ_3 symmetry are forbidden. The absence of emission from the 3P_0 level, which has Γ_1 symmetry, to the first excited Stark level at 19 cm^{-1} indicates that the level has Γ_1 symmetry. Since electric-dipole transitions from the ground state to excited Stark levels having Γ_3 symmetry are forbidden, absorption lines corresponding to hot bands separated by 31 cm^{-1} , but lacking a low-temperature electronic origin, can be used to establish excited Γ_3 Stark levels.

Our analysis of the optical spectra reported in the next section is based on the following points:

- a. Levels observed in the 3P_0 emission spectrum have either Γ_2 , Γ_3 or Γ_4 symmetry; levels established from other measurements, but absent from the 3P_0 emission spectrum, have Γ_1 symmetry.
- b. Levels observed in the 1D_2 emission have either Γ_1 , Γ_3 or Γ_4 symmetry based on the analysis of the hot bands associated with the $16,409\text{ cm}^{-1}$ Stark level, which has Γ_2 symmetry; levels observed in emission spectra from both 3P_0 and 1D_2 have Γ_3 or Γ_4 symmetry; if the emission terminates on a level observed in absorption, then since the ground state has Γ_3 symmetry, the terminal level will have Γ_4 symmetry.
- c. In absorption, if hot-band transitions are present that are 50 cm^{-1} to the low-energy side of a transition that persists at liquid helium temperature, but reveal no 19-cm^{-1} hot band, then the persisting transition is identified as due to an excited Γ_1 Stark level.
- d. Levels observed in the absorption spectrum recorded at 1.6 K have Γ_1 , Γ_2 or Γ_4 symmetry. If transitions to these levels show evidence of hot bands at 19 and 50 cm^{-1} to the low-energy side of the persisting transition, the excited states have Γ_2 symmetry.
- e. Excited Γ_3 Stark levels can be located by observing a pair of hot bands in absorption separated by 31 cm^{-1} and inferring the energy of the forbidden $\Gamma_3 \rightarrow \Gamma_3$ transition.
- f. If only a single 19-cm^{-1} hot band is associated with a transition that persists at liquid helium temperature, the excited Stark level has Γ_4 symmetry.

4. Absorption Spectra

Tables 3 to 5 (given at the end of this section) summarize the results of absorption measurements recorded at 1.6 K, liquid helium temperature (LHe), and liquid nitrogen temperature (LN). The observed spectra appear in figures 1 to 8. From the total number of absorption peaks observed, many of which are weak, it is clear that the Pr^{3+} ions occupy several kinds of sites. The weak peaks retain their relative strengths at low concentrations and hence are not thought to be associated with pairs or clusters of Pr^{3+} ions. It is difficult to determine whether weak peaks are due to weak transitions of Pr^{3+} ions in D_2 sites or strong transitions of Pr^{3+} ions in minority sites. Only the most intense spectra in a given manifold are assigned to transitions of Pr^{3+} ions in D_2 sites. These transitions are indicated in the fourth columns in tables 3 to 5.

Transitions corresponding to the strongest absorption peaks observed are identified using labels introduced by Dieke [21]. More than 100 temperature-dependent (hot-band) transitions were used to establish the first two excited Stark levels, Z_2 at 19 cm^{-1} and Z_3 at 50 cm^{-1} , in agreement with values reported earlier [7] and in agreement with the results obtained from the emission spectra from 3P_0 and 1D_2 . The absence of a transition in absorption and emission between the 19 cm^{-1} (3H_4) and $20,534\text{ cm}^{-1}$ (3P_0) levels indicates that Z_2 (19 cm^{-1}) has Γ_1 symmetry. We chose Γ_4 symmetry for the 50-cm^{-1} level based on our crystal-field splitting calculation; the analysis of the hot-band data is consistent with this choice.

The infrared absorption spectra of the 3H_5 , 3H_6 , and 3F_2 manifolds recorded by the FTIR spectrometer are presented in figures 1 and 2. The absorption peaks become increasingly sharp toward the low-energy side of each manifold since the higher energy levels are broadened by spontaneous phonon emission. This is typical of solid-state rare-earth ion spectra. Multiphonon absorption due to the garnet lattice appears below 2200 cm^{-1} . The spectra recorded at LN temperature are reported in order to show several relatively intense transitions from Z_3 at 50 cm^{-1} . Similar, relatively intense hot bands originating from Z_3 are observed throughout the spectrum.

Figures 3 and 4 display the near-infrared absorption spectra of the 3F_4 , 3F_3 , and 1G_4 manifolds. Proximity between the 3F_4 and 3F_3 manifolds makes it necessary to consider a combined manifold of Stark levels for analysis. The intense sharp spectra and the hot-band assignments

generally agree with the analysis reported by Hooge [7]. The relatively weak 1G_4 absorption spectrum does not contain all the expected transitions. The lack of strong absorption peaks makes it difficult to determine the relative ordering among Stark levels and the overall splitting of the manifold. The predicted crystal-field splitting of the manifold suggests that not all the Stark levels have been observed.

Figure 1. Absorption spectrum of 1H_4 manifold recorded at liquid nitrogen temperature.

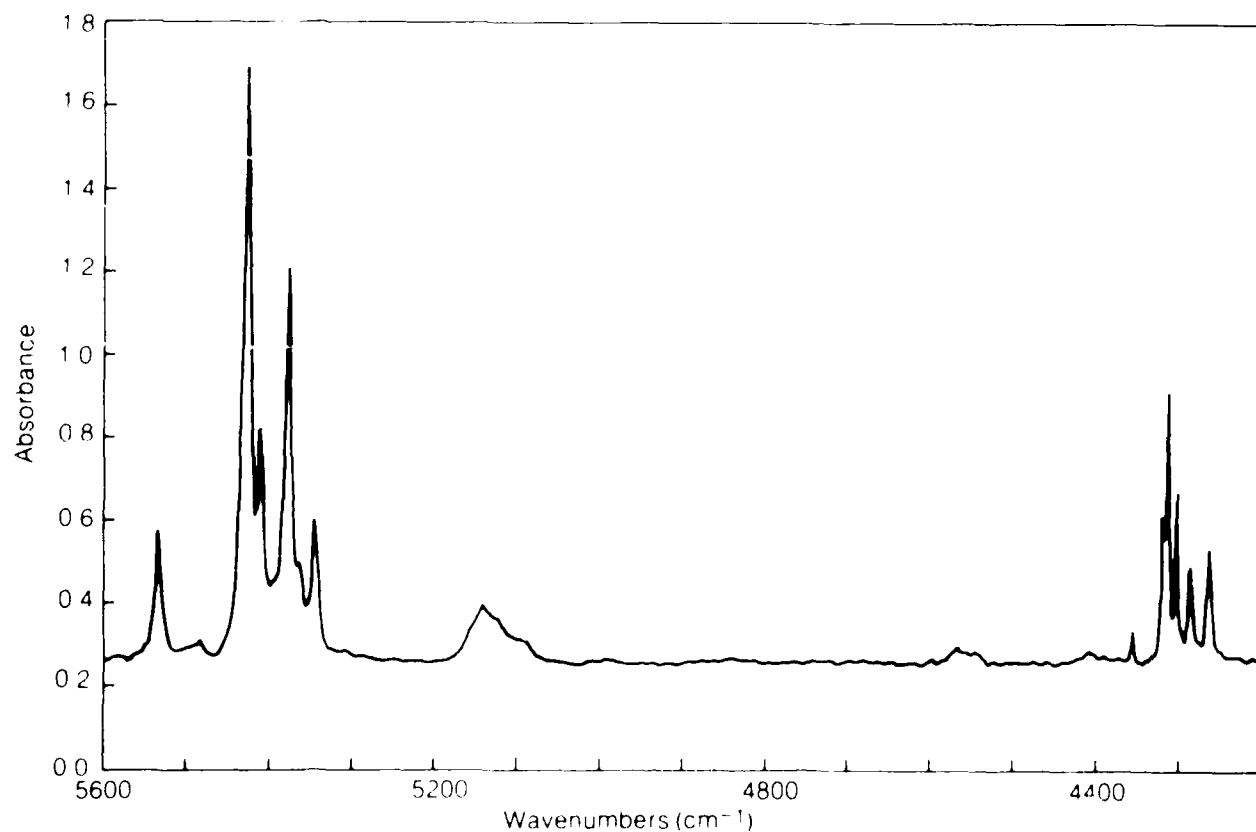
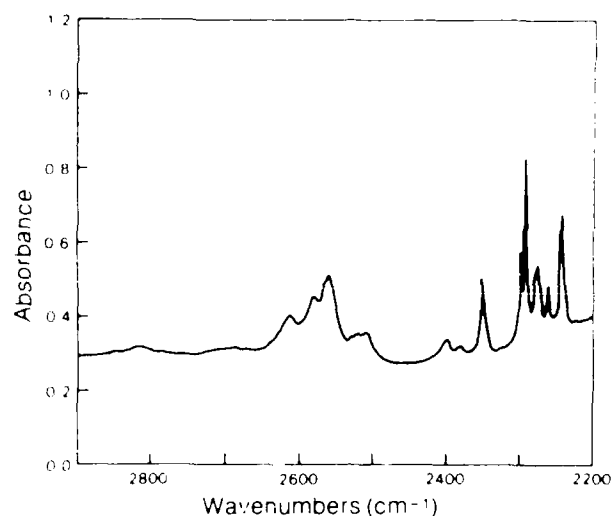


Figure 2. Absorption spectra of 1F_2 and 1H_4 manifolds recorded at 90 K.

To eliminate hot-band absorption entirely, absorption measurements of the 1D_2 , 3P_0 , 3P_1 , 1I_6 and 3P_2 manifolds were made at 1.6 K. The result is the relatively simple spectra appearing in figures 5 to 8. Four principal absorption peaks, representing transitions from Z_1 to B_1 , B_2 , B_3 and B_5 of the 1D_2 manifold appear in figure 5 and are listed in table 5. From the higher temperature data, two hot bands 31 cm^{-1} apart establish the B_4 level at $17,088\text{ cm}^{-1}$. The symmetry of the B_4 level is Γ_3 , following the electric-dipole selection rules of table 2. Two weak peaks at $16,800$ and $16,923\text{ cm}^{-1}$ are assigned to Pr^{3+} ions in other sites.

The 3P_0 absorption at $20,534\text{ cm}^{-1}$ ($Z_1 \rightarrow C_1$) (fig. 6) is assigned to Pr^{3+} ions in D_2 sites in agreement with earlier work [7-11]. The absorption is strong and is comparable to that observed from the ground state to levels W_3 , V_2 , V_5 , B_2 , E_1 , and E_7 . The weak peaks found on either side are attributed to Pr^{3+} ions in other sites. The intense peak at $20,534\text{ cm}^{-1}$ is also observed in fluorescence at 1.6 K. The predicted crystal-field splitting of the 1I_6 manifold (tables 6 and 7) is approximately 1500 cm^{-1}

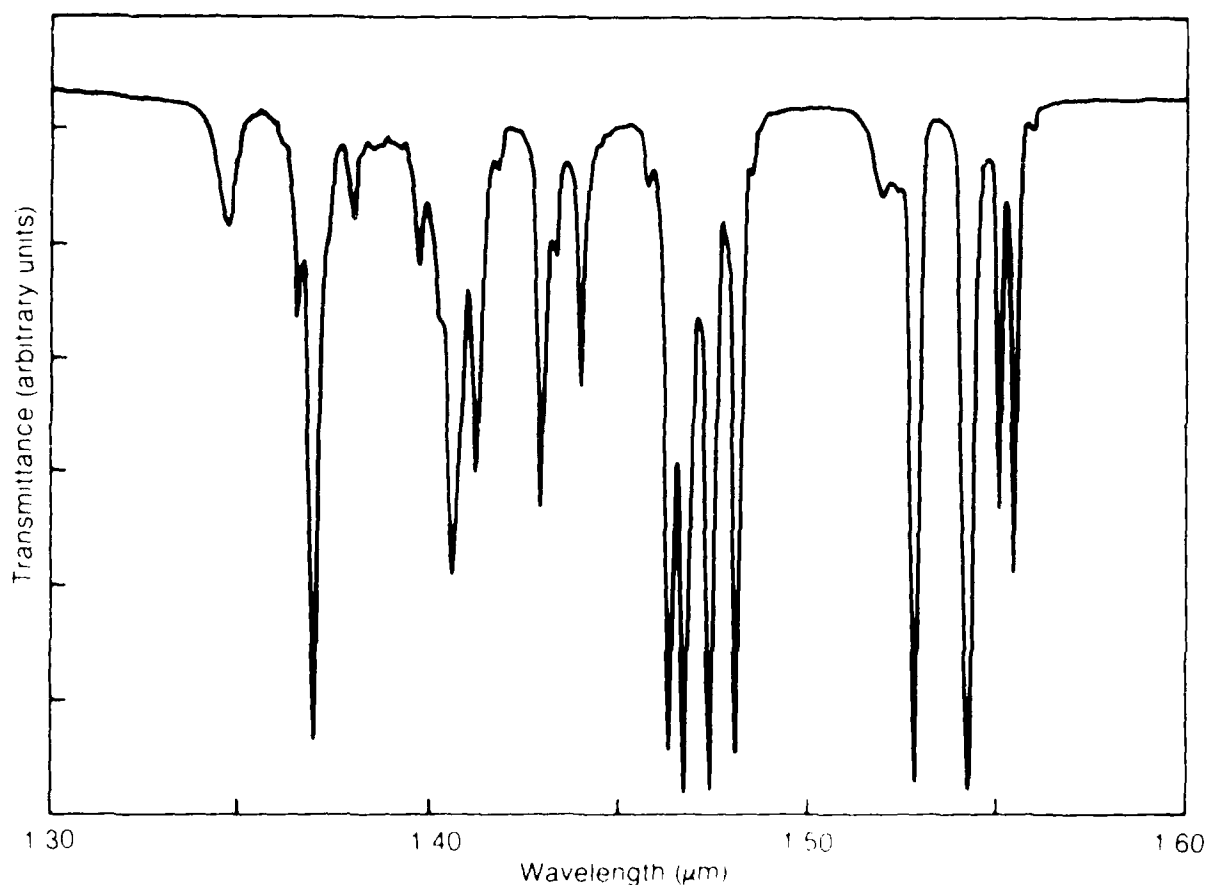


Figure 3. Absorption spectra of 1F_4 and 1F_3 manifolds recorded at 15 K.

and overlaps the predicted splitting of the 3P_1 manifold. We assume that the strong absorption at 1.6 K found at 21,045 and 21,113 cm^{-1} is part of the 3P_1 manifold since $^3H_4 \rightarrow ^1I_6$ absorption is expected to be weak. Figure 7 indicates that only a few transitions to the 1I_6 manifold were observed. The number is insufficient to establish the overall splitting of the 1I_6 manifold. At 1.6 K, absorption peaks were observed at 20,805, 20,823, 21,667, and 21,869 cm^{-1} .

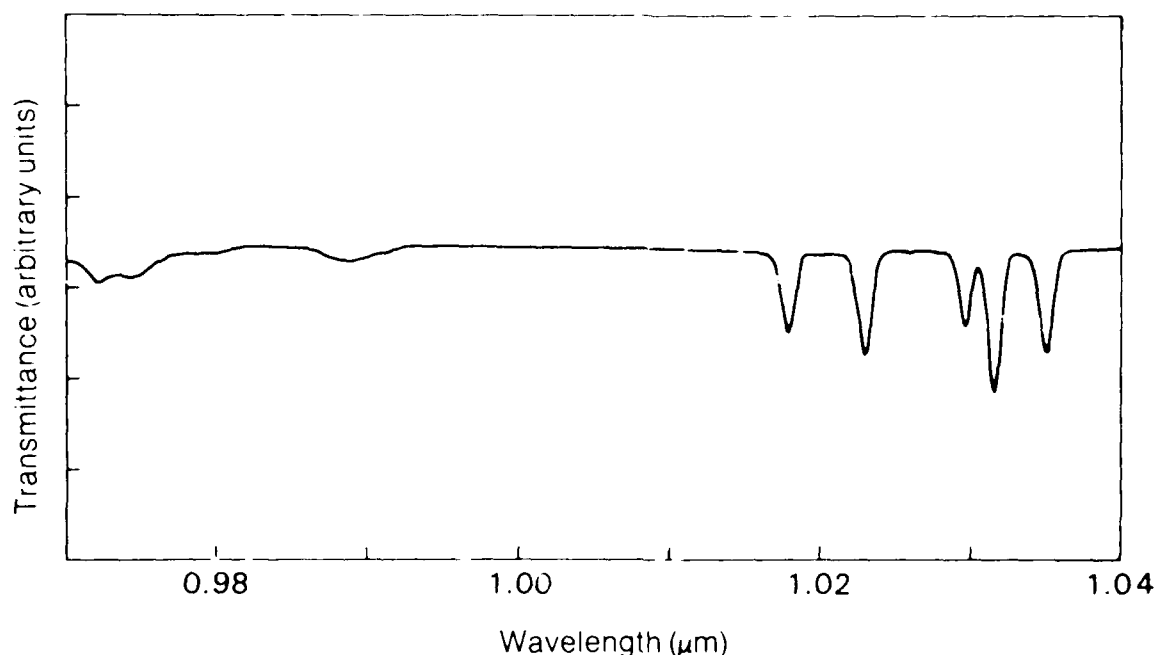


Figure 4. Absorption spectrum of 1G_4 manifold recorded at 15 K.

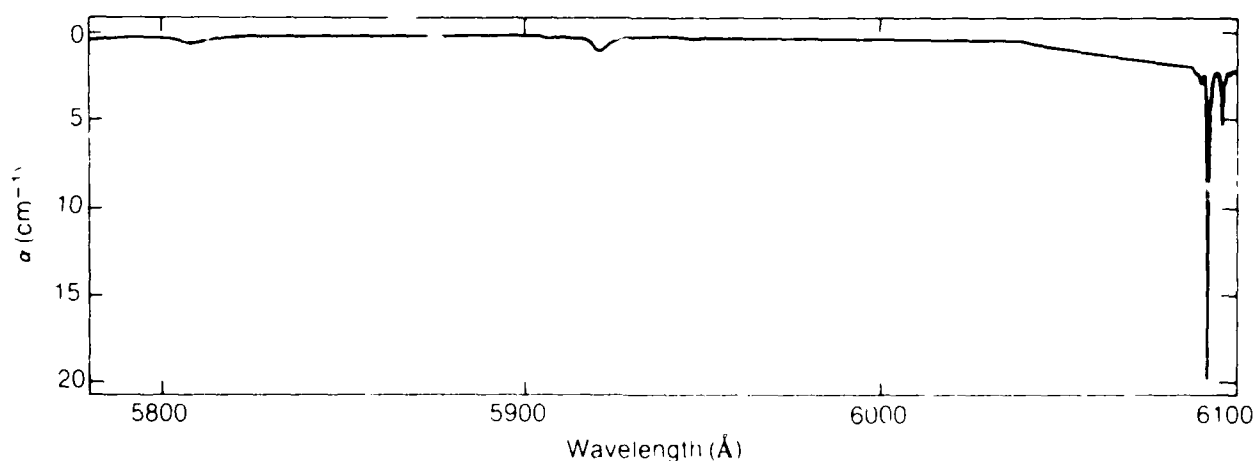


Figure 5. Absorption spectrum of 1D_2 manifold recorded at 1.6 K.

At 1.6 K the absorption to the 3P_2 manifold observed between 22,100 and 22,600 cm^{-1} consists of three peaks. Hot bands assist in making the following assignments: 22,103 (Γ_2), 22,153 (Γ_1), and 22,295 (Γ_1) cm^{-1} . Particularly noticeable even at 1.6 K (fig. 8) are the broad absorption bands that result from spontaneous phonon decay. A pair of hot bands places the Γ_3 level at 22,416 cm^{-1} .

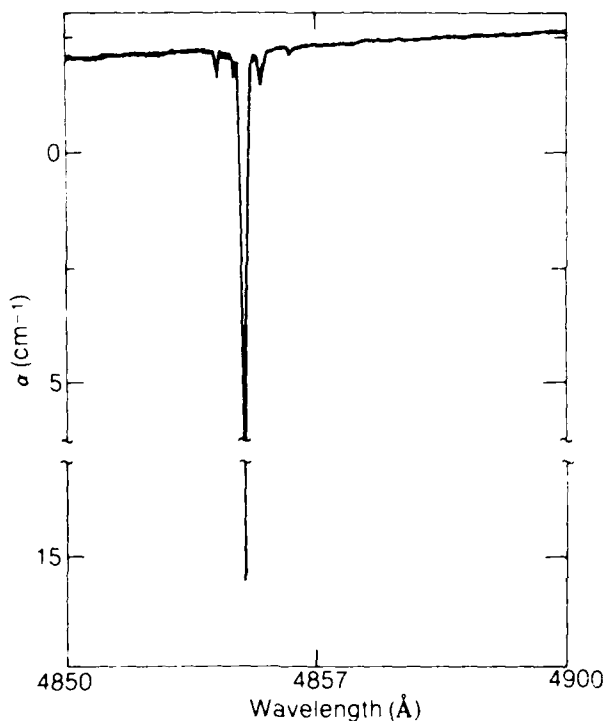


Figure 6. Absorption spectrum of 3P_0 at 1.6 K; four weak, sharp peaks clustered around intense peak suggest that Pr^{3+} ions are found in minority sites.

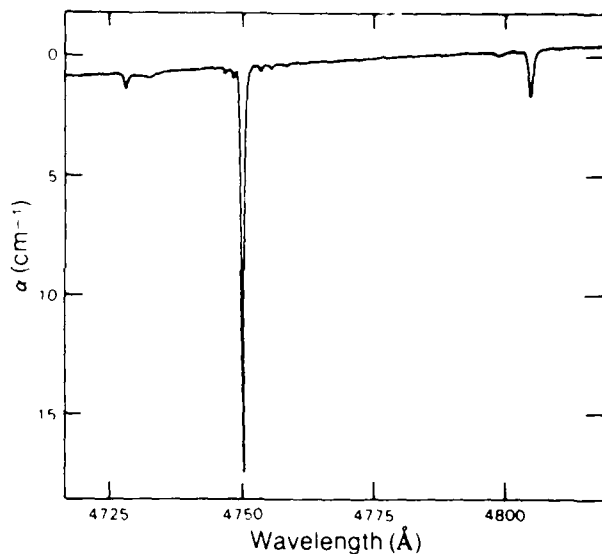


Figure 7. Absorption spectra of 3P_1 and 1I_6 manifolds recorded at 1.6 K.

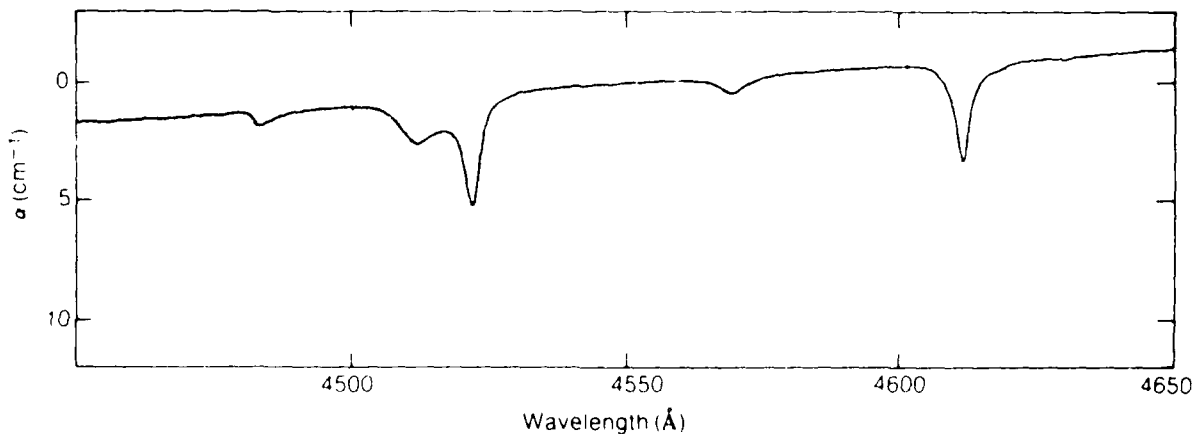


Figure 8. Absorption spectrum of 3P_2 manifold recorded at 1.6 K.

Table 3.
Absorption
and emis-
sion spectra
of 3H_f
manifolds

Note: sh
denotes
shoulder; bd
denotes
broad; ...
indicates that
presence of
level is
inferred (see
sect. 3.
paragraph e).

$2S+1L_J$	E (cm $^{-1}$) ^a	I^b	Transition ^c	$^3P_0^d$ (cm $^{-1}$)	$^1D_2^e$ (cm $^{-1}$)	Level ^f (cm $^{-1}$)	Γ_g^g
3H_4	0	—	(Z_1)	0	0	0	3 (ref. 8)
	19	—	(Z_2)	—	18	19	1 (a, b, c)
	50	—	(Z_3) ^h	51	51	50	4 (a, b, f)
	—	—	(Z_4)	—	533	533	1 (a, b)
	—	—	(Z_5)	576	—	576	2-4 (a)
	—	—	(Z_6)	—	742	742	1 (a, b)
3H_5	2209	<0.01	—	—	—	—	—
	2221	0.01	—	—	—	—	—
	2242*	0.27	$Z_3 \rightarrow Y_1$	—	—	—	—
	2261*	0.17	$Z_3 \rightarrow Y_3$	—	—	—	—
	2274*	0.20	$Z_2 \rightarrow Y_1$	—	—	—	—
	2280*	0.17	$Z_2 \rightarrow Y_2$	—	—	—	—
	2293	0.41	$Z_1 \rightarrow Y_1$	—	—	2293	2 (d)
	2299	0.36	$Z_1 \rightarrow Y_2$	2295	2295	2299	4 (a,b,f)
	(Y_3)	—	—	(2311)	3 (e)
	2347*(sh)	0.05	—	—	—	—	—
	2351	0.14	$Z_1 \rightarrow Y_4$	—	—	2351	1 (c)
	2380*	0.06	$Z_2 \rightarrow Y_5$	—	—	—	—
	2399	0.09	$Z_1 \rightarrow Y_5$	2398	2398	2399	(a,d)
	2520*	0.09	$Z_3 \rightarrow Y_6$	—	—	—	—
	2551*	0.10	$Z_2 \rightarrow Y_6$	—	—	—	—
	(Y_6)	—	—	(2570)	3 (e)
	2562*	0.20	$Z_2 \rightarrow Y_7$	—	—	—	—
	2581	0.18	$Z_1 \rightarrow Y_7$	2580	2580	2581	4 (a,b,f)
	2599	0.06	$Z_2 \rightarrow Y_9$	—	—	—	—
	2602	0.12	$Z_1 \rightarrow Y_8$	—	—	2602	1 (c)
	2618	0.10	$Z_1 \rightarrow Y_9$	—	—	2618	2 (d)
	2669	bd	—	—	—	—	—
	2688	bd	—	—	—	—	—
	2752	0.01	—	—	—	—	—
	2790*	0.02	$Z_3 \rightarrow Y_{11}$	—	—	—	—
	2800*	0.03	$Z_2 \rightarrow Y_{10}$	—	—	—	—
	2820	0.04	$Z_1 \rightarrow Y_{10}$	—	—	2820	4 (f)
	2822*(sh)	0.02	$Z_2 \rightarrow Y_{11}$	—	—	—	—
	(Y_{11})	2841	2841	2840	3 (a,b,e)
	2850	bd	—	—	—	—	—
3H_6	4214	0.01	—	—	—	—	—
	4244	0.02	—	—	—	—	—
	4265*	0.11	$Z_3 \rightarrow X_2$	—	—	—	—
	4286*	0.11	$Z_2 \rightarrow X_1$	—	—	—	—
	4289*	0.12	$Z_3 \rightarrow X_3$	—	—	—	—
	4305	0.24	$Z_1 \rightarrow X_1$	—	—	4305	4 (f)
	4316	0.35	$Z_1 \rightarrow X_2$	—	—	4316	1 (c)

Table 3
(cont'd).
Absorption
and emission
spectra of 3H_4
manifolds

$^{25+1}L_J$	E (cm $^{-1}$) ^a	I ^b	Transition ^c	3P_0 ^d (cm $^{-1}$)	1D_2 ^e (cm $^{-1}$)	Level ^f (cm $^{-1}$)	Γ_4 ^g
3H_6	4321*	0.17	$Z_2 \rightarrow X_3$	—	—	—	—
(cont'd)	(X_3)	—	—	4339	3 (e)
	4347	0.01	—	—	—	—	—
	4358	0.06	$Z_1 \rightarrow X_4$	—	—	4358	1 (c)
	4391*	0.02	$Z_2 \rightarrow X_5$	—	—	—	—
	4403*(sh)	0.02	—	—	—	—	—
	4409	0.03	$Z_1 \rightarrow X_5$	—	—	4409	2 (d)
	4420*	0.02	$Z_2 \rightarrow X_6$	—	—	—	—
	(X_6)	—	—	4440	3 (e)
	4548*	0.03	$Z_2 \rightarrow X_7$	—	—	—	—
	4558*	0.02	—	—	—	—	—
	4567	0.04	$Z_1 \rightarrow X_7$	—	—	4567	4 (f)
	4793*	0.01	$Z_3 \rightarrow X_{11}$	—	—	—	—
	4824*	0.01	$Z_2 \rightarrow X_{11}$	—	—	—	—
	4843	0.02	$Z_1 \rightarrow X_{11}$	—	—	4843	2 (d)
	4994	0.02	$Z_3 \rightarrow X_{12}$	—	—	—	—
	5045	0.03	$Z_1 \rightarrow X_{12}$	—	—	5045	1 (c)
	5091	0.03	$Z_3 \rightarrow X_{13}$	—	—	—	—
	5122*	0.10	$Z_2 \rightarrow X_{13}$	—	—	—	—
	5141	0.12	$Z_1 \rightarrow X_{13}$	—	—	5141	2 (d)

^a Absorption spectra recorded at LHe temperature for 3H_6 , LN temperature for 3H_5 ; room temperature spectra for both manifolds were used to establish hot bands, denoted by *.

^b Intensity in absorbance units.

^c Transitions assigned to absorption spectra of Pr^{3+} in D_2 sites.

^d Fluorescence spectrum from 3P_0 at 1.6 K.

^e Selectively pumped fluorescence spectrum from 1D_2 at 1.6 K.

^f Experimental energy levels assigned to Pr^{3+} in D_2 sites.

^g D_2 symmetry representations Γ_1 , Γ_2 , Γ_3 , and Γ_4 ; letters in parentheses refer to method of assignment described in sect. 3.

^h Ref. 8 identifies Z_3 as either Γ_2 or Γ_4 ; observed emission from the 16,409 cm $^{-1}$ level (which has Γ_2 symmetry based on an analysis of hot bands in the absorption spectrum) indicates that Z_3 has Γ_4 symmetry, in agreement with calculations in table 7.

Table 4. Absorption spectra of 3F_1 manifolds

Note: sh denotes shoulder; bd denotes broad; . . . indicates that presence of level is inferred (see sect. 3, paragraph e).

$2S+1L_J$	E (cm $^{-1}$) ^a	I^b	Transition ^c	Level ^d (cm $^{-1}$)	Γ_e
3F_2	5290	<0.01	—	—	—
	5305	0.01	—	—	—
	5346*	0.32	$Z_2 \rightarrow W_1$	—	—
	5365	0.23	$Z_1 \rightarrow W_1$	5365	4 (f)
	5379*	0.92	$Z_3 \rightarrow W_2$	—	—
	5411*	0.55	$Z_2 \rightarrow W_2$	—	—
	5430	1.42	$Z_1 \rightarrow W_2$	5430	2 (d)
	5485*	0.05	$Z_3 \rightarrow W_4$	—	—
	5516*	0.04	$Z_2 \rightarrow W_4$	—	—
	(W_4)	(5535)	3 (c)
	5535*	0.32	$Z_3 \rightarrow W_5$	—	—
	5585	0.03	$Z_1 \rightarrow W_5$	5585	1 (c)
	5621	<0.01	—	—	—
3F_3	6417*	0.20	$Z_3 \rightarrow V_1$	—	—
	6432*	0.70	$Z_3 \rightarrow V_2$	—	—
	6448*	0.86	$Z_2 \rightarrow V_1$	—	—
	(V_1)	(6467)	3 (c)
	6482	1.30	$Z_1 \rightarrow V_2$	6482	1 (c)
	6511	0.15	—	—	—
	6541*	1.30	$Z_2 \rightarrow V_3$	—	—
	6561	0.40	$Z_1 \rightarrow V_3$	6561	4 (f)
	6572	0.03	—	—	—
	6590	0.03	—	—	—
	6760	0.90	$Z_2 \rightarrow V_5$	—	—
	6779*(sh)	0.80	$Z_3 \rightarrow V_6$	—	—
	6781	1.40	$Z_1 \rightarrow V_5$	6781	4 (f)
	6812*	1.40	$Z_2 \rightarrow V_6$	—	—
	6831(sh)	1.12	$Z_1 \rightarrow V_6$	6831	2 (d)
	6857	0.02	—	—	—
	6876	0.02	—	—	—
	6943*	0.20	$Z_3 \rightarrow V_7$	—	—
	6973*	0.10	$Z_2 \rightarrow V_7$	—	—
	6994	0.30	$Z_1 \rightarrow V_7$	6994	2 (d)
3F_4	7035*	0.06	$Z_3 \rightarrow U_3$	—	—
	7041*	0.05	$Z_2 \rightarrow U_2$	—	—
	7060	0.03	$Z_1 \rightarrow U_2$	7060	4 (f)
	7085	0.40	$Z_1 \rightarrow U_3$	7085	1 (c)
	7092*	0.30	$Z_3 \rightarrow U_4$	—	—
	7122*	0.65	$Z_2 \rightarrow U_4$	—	—
	7142	0.30	$Z_1 \rightarrow U_4$	7142	2 (d)
	7161	0.03	—	—	—
	7194*	0.05	$Z_3 \rightarrow U_5$	—	—

Table 4 (cont'd).
Absorption
spectra of 3F_4
manifolds

$^{25+1}L_J$	E (cm $^{-1}$) ^a	I^b	Transition ^c	Level ^d (cm $^{-1}$)	Γ_e^e
3F_4 (cont'd)	7245	0.09	$Z_1 \rightarrow U_5$	7245	1 (c)
	7280*	0.06	$Z_2 \rightarrow U_6$	—	—
	7297	0.75	$Z_1 \rightarrow U_6$	7297	4 (f)
	7319	0.25	$Z_1 \rightarrow U_7$	7319	2 (d)
	7370*	0.05	$Z_3 \rightarrow U_8$	—	—
	7402*	0.10	$Z_2 \rightarrow U_8$	—	—
	(U_9)	(7421)	3 (e)
	7440*	0.10	$Z_3 \rightarrow U_9$	—	—
	7490	0.05	$Z_1 \rightarrow U_9$	7490	1 (c)

^a Spectra recorded on spectrophotometer at LHe temperature.

^b Intensity in absorbance units.

^c Transitions assigned to spectra of Pr^{3+} in D_2 sites.

^d Experimental energy levels assigned to Pr^{3+} in D_2 sites.

^e D_2 symmetry representations given to assigned Stark levels; letters in parentheses refer to method of assignment described in sect. 3.

Table 5.
Absorption
spectra of 1G_4 ,
 1D_2 , 3P_0 , 1I_1 , 3P_1 ,
and 3P_2
manifolds

Note: * refers to
observed hot
bands at LN
temperature; sh
denotes shoul-
der; bd = broad,
usually very
weak, Pr^{3+} site
uncertain; ...
indicates that
presence of level
is inferred (see
sect. 3, para-
graph e).

$^{2S+1}L_J$	$E \text{ (cm}^{-1}\text{)}^a$	f^b	Transition ^c	$E \text{ (cm}^{-1}\text{)}^d$	Level ^e (cm^{-1})	Γ_s^f
1G_4	9,667*	0.08	$Z_3 \rightarrow A_1$	—	—	—
	9,698*	0.12	$Z_2 \rightarrow A_1$	—	—	—
	9,717	0.06	$Z_1 \rightarrow A_1$	—	9,717	2 (d)
	9,775*	bd	—	—	—	—
	9,780*	0.10	$Z_3 \rightarrow A_4$	—	—	—
	9,828	0.07	$Z_1 \rightarrow A_4$	—	9,828	1 (c)
	10,096*	0.02	$Z_2 \rightarrow A_6$	—	—	—
	10,116	0.03	$Z_1 \rightarrow A_6$	—	10,116	4 (f)
	10,211*	0.01	$Z_3 \rightarrow A_7$	—	—	—
	10,262*	0.03	$Z_2 \rightarrow A_7$	—	—	—
	10,282	0.03	$Z_1 \rightarrow A_7$	—	10,282	2 (d)
	10,330*	0.02	$Z_3 \rightarrow A_8$	—	—	—
	10,360*	0.02	$Z_2 \rightarrow A_8$	—	—	—
	(A_8)	—	(10,380)	3 (e)
1D_2	16,264	0.01	—	—	—	—
	16,324	0.02	—	—	—	—
	16,350*(sh)	0.30	$Z_3 \rightarrow B_1$	—	—	—
	16,358*	0.83	$Z_3 \rightarrow B_2$	—	—	—
	16,381(sh)	0.03	—	—	—	—
	16,390*	0.50	$Z_2 \rightarrow B_2$	—	—	—
	16,400	0.46	$Z_1 \rightarrow B_1$	16,400	16,400	1 (c)
	16,409	1.10	$Z_1 \rightarrow B_2$	16,410	16,409	2 (b,d)
	16,478	0.01	—	—	—	—
	16,755	0.01	—	—	—	—
	16,777	<0.01	—	—	—	—
	16,805	0.05	—	16,800	—	—
	16,830*	1.19	$Z_3 \rightarrow B_3$	—	—	—
	16,856	0.03	—	—	—	—
	16,881	0.33	$Z_1 \rightarrow B_3$	16,887	16,881	1 (c)
	16,903	0.03	—	—	—	—
	16,921	0.03	—	16,919	—	—
	16,956	<0.01	—	—	—	—
	17,037*	0.26	$Z_3 \rightarrow B_4$	—	—	—
	17,069*	0.06	$Z_2 \rightarrow B_4$	—	—	—
	(B_4)	—	(17,088)	3 (e)
	17,191(sh)*	0.07	$Z_2 \rightarrow B_5$	—	—	—
	17,210	0.17	$Z_1 \rightarrow B_5$	17,207	17,210	4 (f)
	17,250	bd	—	—	—	—
3P_0	20,459	0.04	—	—	—	—
	20,483*	0.20	$Z_3 \rightarrow C_1$	—	—	—
	20,517	0.05	—	—	—	—
	20,534	0.60	$Z_1 \rightarrow C_1$	20,530	20,534	1 (a,c)

Table 5
(cont'd).
Absorption
spectra of 1G_4 ,
 1D_2 , 3P_0 , 1I_6 , 3P_1 ,
and 3P_2
manifolds

$^{2S+1}L_J$	E (cm $^{-1}$) ^a	I ^b	Transition ^c	E (cm $^{-1}$) ^d	Level ^e (cm $^{-1}$)	Γ_s ^f
1I_6	20,681	0.02	—	—	—	—
and	20,701	0.03	—	—	—	—
3P_1	20,755*	0.03	—	—	—	—
	20,781*	0.03	—	—	—	—
	20,786*	0.10	—	—	—	—
	20,805	0.20	—	20,805	—	—
	20,831	0.05	—	20,823	—	—
	20,942	<0.01	—	—	—	—
	20,950	<0.01	—	—	—	—
	20,994*	0.51	$Z_3 \rightarrow E_1$	—	—	—
	21,010*(sh)	0.04	—	—	—	—
	21,026*	0.23	$Z_2 \rightarrow E_1$	—	—	—
	21,045	1.32	$Z_1 \rightarrow E_1$	21,045	21,045	2 (d , 3P_1)
	21,087*	0.09	—	—	—	—
	21,093*	0.09	$Z_2 \rightarrow E_2$	—	—	—
	21,113	0.87	$Z_1 \rightarrow E_2$	21,113	21,113	4 (f , 3P_1)
	21,140	0.03 ^b	—	21,135	—	—
	21,152*	0.25	—	—	—	—
	21,520*	0.03	—	—	—	—
	21,568*	0.05	—	—	—	—
	...	—	—	—	—	—
	21,588	0.15	—	—	—	—
	21,651*	0.60	—	—	—	—
	21,672	1.10	—	21,667	—	—
	21,732	0.03	—	—	—	—
	21,781	0.01	—	—	—	—
	21,810	0.01	—	—	—	—
	21,820*	0.10	—	—	—	—
	21,870	0.27	—	21,869	—	—
3P_2	22,054*	0.50	$Z_3 \rightarrow F_1$	—	—	—
	22,084*	0.95	$Z_2 \rightarrow F_1$	—	—	—
	22,103	1.12	$Z_1 \rightarrow F_1$	22,100	22,103	2 (d)
	22,153	0.60	$Z_1 \rightarrow F_2$	22,150	22,153	1 (c)
	22,245*	0.14	$Z_3 \rightarrow F_3$	—	—	—
	22,295	0.25	$Z_1 \rightarrow F_3$	22,288	22,295	1 (c)
	22,367	0.08	$Z_1 \rightarrow F_4$	—	22,367	4 (f)
	22,397	0.12	$Z_2 \rightarrow F_5$	—	—	—
	(F_5)	—	(22,416)	3 (d)
	22,570	bd	—	—	—	—

^a Spectra recorded on spectrophotometer at LHe temperature.

^b Intensity in absorbance units.

^c Transitions assigned to spectra of Pr^{3+} in D_2 sites.

^d Spectra recorded at 1.6 K.

^e Observed energy levels of Pr^{3+} in D_2 sites.

^f D_2 symmetry representations given to assigned Stark levels; letters in parentheses refer to method of assignment given in sect. 3.

Table 6. Cubic O_h^{10} ($1a3d$) 230

Lattice constant is 1.2 nm (ref. 18); crystallographic data on $Y_3Al_5O_{12}$

Ion	Site	Symmetry	x	y	z	q^a	$\alpha (\text{\AA}^3)^b$
Y	24c	D_2	0	1/4	1/8	3	0.870
Al ₁	16a	C_{3v}	0	0	0	3	0.0530
Al ₂	24d	S_6	0	1/4	3/8	q_{Al2}	0.0530
O	96h	C_1	x^c	y^c	z^c	q_o	1.349

^a The values of q are in units of the electron charge with $4q_o + q_{Al2} = -5$.

^b Ref. 26.

^c $x = -0.0306$, $y = 0.0512$, $z = 0.150$ (ref. 18).

Table 7. Energy levels of Pr^{3+} ions in D_2 sites

Note: ... indicates that presence of level is inferred (see sect. 3, paragraph e)

$2s+1L_J$	Experimental ^a		Calculated ^b		Percent free-ion mixture ^c
	$E (\text{cm}^{-1})$	Γ_n	$E (\text{cm}^{-1})$	Γ_n	
3H_4	0	3	-1	3	$99.9\ ^3H_4 + 0.08\ ^3H_5 + 0.03\ ^3F_3$
503	19	1	19	1	$99.9\ ^3H_4 + 0.05\ ^3F_2 + 0.04\ ^3H_5$
	50	4 ^c	39	4	$99.6\ ^3H_4 + 0.21\ ^3H_5 + 0.08\ ^3F_2$
	495	2	$95.2\ ^3H_4 + 3.55\ ^3H_5 + 1.03\ ^3F_2$
	533	1 ^c	505	1	$98.1\ ^3H_4 + 0.77\ ^3F_2 + 0.43\ ^3F_4$
	506	3	$96.1\ ^3H_4 + 2.90\ ^3H_5 + 0.75\ ^3F_3$
	576	4 ^c	532	4	$95.6\ ^3H_4 + 3.39\ ^3H_5 + 0.88\ ^3F_3$
	742	1	755	1	$93.6\ ^3H_4 + 5.17\ ^3H_5 + 0.91\ ^3F_2$
	759	2	$93.3\ ^3H_4 + 5.75\ ^3H_5 + 0.54\ ^3F_3$
3H_5	2293	2	2286	2	$96.8\ ^3H_5 + 2.26\ ^3H_4 + 0.79\ ^3F_2$
2595	2299	4	2270	4	$95.5\ ^3H_5 + 3.28\ ^3H_4 + 0.42\ ^3F_3$
	2311	3 ^c	2291	3	$97.3\ ^3H_5 + 1.34\ ^3H_4 + 1.09\ ^3F_3$
	2351	1	2357	1	$93.3\ ^3H_5 + 4.28\ ^3H_4 + 1.06\ ^3F_2$
	2399	2	2412	2	$95.1\ ^3H_5 + 3.59\ ^3H_4 + 0.45\ ^3H_6$
	2570	3	2577	3	$93.6\ ^3H_5 + 3.10\ ^3F_2 + 2.05\ ^3H_6$
	2581	4	2583	4	$94.4\ ^3H_5 + 2.73\ ^3H_6 + 1.06\ ^3F_2$
	2602	1	2597	1	$93.4\ ^3H_5 + 3.92\ ^3F_2 + 1.62\ ^3H_6$
	2618	2 ^f	2625	2	$94.3\ ^3H_5 + 2.53\ ^3H_6 + 1.44\ ^3H_4$
	2820	4	2824	4	$92.4\ ^3H_5 + 4.71\ ^3H_6 + 2.25\ ^3F_2$
	2840	3	2836	3	$92.6\ ^3H_5 + 5.13\ ^3H_6 + 0.79\ ^3F_2$
3H_6	4305	4	4302	4	$94.2\ ^3H_6 + 4.17\ ^3F_2 + 0.90\ ^3H_5$
4744	4316	1	4325	1	$95.9\ ^3H_6 + 2.97\ ^3F_2 + 0.49\ ^3H_5$
	4339	3	4365	3	$95.7\ ^3H_6 + 3.22\ ^3F_2 + 0.61\ ^3F_4$
	4358	1	4347	1	$95.3\ ^3H_6 + 3.37\ ^3F_2 + 0.60\ ^3F_4$
	4409	2	4398	2	$95.6\ ^3H_6 + 1.84\ ^3H_5 + 1.57\ ^3F_4$
	4440	3	4431	3	$93.6\ ^3H_6 + 3.89\ ^3H_5 + 0.96\ ^3F_4$
	4567	4	4574	4	$92.1\ ^3H_6 + 3.78\ ^3H_5 + 2.09\ ^3F_4$
	4695	1	$88.1\ ^3H_6 + 5.80\ ^3F_4 + 4.77\ ^3F_2$
	4736	3	$75.4\ ^3H_6 + 17.0\ ^3F_2 + 4.12\ ^3F_3$
	4780	4	$78.1\ ^3H_6 + 12.6\ ^3F_2 + 7.66\ ^3F_3$
	4843	2	4843	2	$86.8\ ^3H_6 + 7.99\ ^3F_2 + 2.53\ ^3F_3$
	5045	1	5047	1	$66.2\ ^3H_6 + 25.7\ ^3F_2 + 6.80\ ^3F_3$
	5141	2 ^f	5106	2	$75.7\ ^3H_6 + 11.8\ ^3F_2 + 9.61\ ^3F_3$

Table 7 (cont'd).
Energy levels of
Pr³⁺ ions in D₂ sites

^{2S+1} L _J	Experimental ^a		Calculated ^b		Percent free-ion mixture ^c
	E (cm ⁻¹)	Γ _s	E (cm ⁻¹)	Γ _s	
³ F ₂ 5292	5365	4	5388	4	74.0 ³ F ₂ + 15.6 ³ H ₆ + 5.21 ³ F ₄
	5422	1	69.1 ³ F ₂ + 22.9 ³ H ₆ + 5.23 ³ F ₄
	5430	2	5424	2	83.4 ³ F ₂ + 14.6 ³ H ₆ + 1.14 ³ H ₄
	5535	3	5530	3	67.9 ³ F ₂ + 17.9 ³ H ₆ + 6.19 ³ F ₄
	5585	1	5575	1	83.1 ³ F ₂ + 11.0 ³ H ₆ + 2.64 ³ H ₄
³ F ₃ 6604	6467	3	6470	3	78.3 ³ F ₃ + 14.3 ³ F ₄ + 5.61 ³ F ₂
	6482	1	6486	1	75.8 ³ F ₃ + 18.7 ³ F ₄ + 3.27 ³ F ₂
	6561	4	6577	1	71.0 ³ F ₃ + 22.4 ³ F ₄ + 3.35 ³ F ₂
	6761	3	95.7 ³ F ₃ + 1.25 ³ H ₄ + 1.18 ³ H ₆
	6781	4	6768	4	92.2 ³ F ₃ + 4.50 ³ F ₄ + 1.15 ³ H ₄
	6831	2'	6803	2	96.0 ³ F ₃ + 1.32 ³ H ₄ + 0.92 ³ H ₆
	6994	2	6999	2	54.6 ³ F ₃ + 32.4 ³ F ₄ + 12.3 ³ H ₆
³ F ₄ 7111	7040	3	78.0 ³ F ₄ + 14.2 ³ F ₃ + 6.25 ³ H ₆
	7060	4'	7050	4	74.1 ³ F ₄ + 18.1 ³ F ₃ + 6.41 ³ H ₆
	7085	1	7088	1	73.1 ³ F ₄ + 16.6 ³ F ₃ + 8.10 ³ H ₆
	7142	2	7149	2	60.4 ³ F ₄ + 28.0 ³ F ₃ + 9.01 ³ H ₆
	7245	1	7250	1	93.5 ³ F ₄ + 4.24 ³ H ₆ + 1.07 ¹ G ₄
	7297	4	7303	4	88.8 ³ F ₄ + 7.13 ³ F ₃ + 2.77 ³ H ₆
	7319	2	7314	2	97.3 ³ F ₄ + 1.20 ³ H ₆ + 0.87 ³ H ₄
	7421	3	7404	3	94.5 ³ F ₄ + 2.35 ³ H ₆ + 1.42 ³ F ₃
	7490	1	7488	1	95.3 ³ F ₄ + 1.90 ¹ G ₄ + 1.69 ³ H ₆
¹ G ₄ 10,042	9,693	3	98.2 ¹ G ₄ + 0.76 ³ H ₆ + 0.62 ³ F ₄
	9,716	1	98.2 ¹ G ₄ + 0.63 ³ H ₆ + 0.57 ³ F ₄
	9,717	2'	9,721	2	97.7 ¹ G ₄ + 1.47 ³ F ₄ + 0.58 ³ H ₆
	9,724	4	98.7 ¹ G ₄ + 0.46 ³ H ₆ + 0.32 ³ F ₄
	9,828	1'	9,907	1	98.3 ¹ G ₄ + 0.78 ³ H ₆ + 0.70 ³ F ₄
	10,116	4	10,118	4	99.0 ¹ G ₄ + 0.32 ¹ I ₆ + 0.25 ³ H ₄
	10,282	2	10,281	2	99.0 ¹ G ₄ + 0.46 ¹ I ₆ + 0.20 ³ H ₄
	10,380	3'	10,380	3	99.0 ¹ G ₄ + 0.36 ¹ I ₆ + 0.26 ³ H ₄
	10,916	1	96.8 ¹ G ₄ + 1.82 ³ F ₄ + 1.19 ¹ I ₆
¹ D ₂ 16,883	16,400	1	16,411	1	99.7 ¹ D ₂ + 0.12 ³ P ₂ + 0.08 ¹ I ₆
	16,409	2	16,417	2	99.5 ¹ D ₂ + 0.17 ³ P ₂ + 0.14 ¹ I ₆
	16,881	1	16,867	1	97.1 ¹ D ₂ + 2.77 ¹ I ₆ + 0.05 ¹ G ₄
	17,088	3	17,082	3	97.6 ¹ D ₂ + 2.31 ¹ I ₆ + 0.05 ¹ G ₄
	17,210	4'	17,117	4	97.4 ¹ D ₂ + 2.28 ¹ I ₆ + 0.12 ¹ G ₄
³ P ₀	20,534	1	20,534	1	83.0 ³ P ₀ + 15.9 ¹ I ₆ + 0.73 ³ P ₂
¹ I ₆	20,555	1	84.1 ¹ I ₆ + 15.6 ³ P ₀ + 0.14 ³ P ₂
	20,560	3	99.8 ¹ I ₆ + 0.07 ¹ G ₄ + 0.07 ³ P ₂
	20,681	4	99.9 ¹ I ₆ + 0.08 ³ P ₂ + 0.02 ¹ D ₂
	20,697	2	99.9 ¹ I ₆ + 0.07 ³ P ₂ + 0.03 ¹ D ₂
	20,937	2	99.9 ¹ I ₆ + 0.04 ³ P ₁ + 0.02 ³ P ₂
	20,938	1	99.9 ¹ I ₆ + 0.02 ³ P ₂ + 0.01 ¹ G ₄

Table 7 (cont'd).
Energy levels of
Pr³⁺ ions in D₂ sites

²⁵⁺¹ L _j	Experimental ^a		Calculated ^b		Percent free-ion mixture ^c
	E (cm ⁻¹)	Γ _a	E (cm ⁻¹)	Γ _a	
³ P ₁	21,045	2	21,056	2	98.5 ³ P ₁ + 1.13 ³ P ₂ + 0.14 ³ F ₃
21,099	21,113	4	21,103	4	97.0 ³ P ₁ + 2.27 ³ P ₂ + 0.40 ¹ I ₆
	21,237	3	99.3 ³ P ₁ + 0.31 ³ P ₂ + 0.09 ¹ I ₆
¹ I ₆	21,456	1	87.7 ¹ I ₆ + 9.93 ³ P ₂ + 2.11 ¹ D ₂
21,250	21,527	3	96.0 ¹ I ₆ + 2.82 ³ P ₂ + 0.95 ¹ D ₂
	21,716	4	84.3 ¹ I ₆ + 13.1 ³ P ₂ + 1.38 ¹ D ₂
	21,834	3	88.9 ¹ I ₆ + 9.79 ³ P ₂ + 0.92 ¹ D ₂
	21,858	4	95.7 ¹ I ₆ + 3.43 ³ P ₂ + 0.38 ³ P ₁
	21,876	2	95.5 ¹ I ₆ + 3.82 ³ P ₂ + 0.42 ¹ G ₄
	22,005	1	75.0 ¹ I ₆ + 23.3 ³ P ₂ + 0.73 ¹ G ₄
³ P ₂	22,103	2	22,076	2	94.6 ³ P ₂ + 3.92 ¹ I ₆ + 1.12 ³ P ₁
22,147	22,153	1	22,172	1	84.9 ³ P ₂ + 14.1 ¹ I ₆ + 0.38 ³ P ₀
	22,295	1	22,317	1	80.4 ³ P ₂ + 18.8 ¹ I ₆ + 0.48 ¹ D ₂
	22,367	4	22,364	4	80.8 ³ P ₂ + 17.0 ¹ I ₆ + 1.29 ³ P ₁
	22,416	3	22,410	3	86.8 ³ P ₂ + 12.4 ¹ I ₆ + 0.43 ¹ D ₂
¹ S ₀	46,950	1	99.8 ¹ S ₀ + 0.08 ¹ I ₆ + 0.06 ¹ G ₄
46,900					

^a Experimental energy levels and symmetry labels established from spectra reported in tables 3 to 5.

^b Theoretical energy levels calculated using B_{km} (cm⁻¹): B₂₀ (481), B₂₂ (123), B₄₀ (-146), B₄₂ (-2248), B₄₄ (-1139), B₆₀ (-1653), B₆₂ (-772), B₆₄ (869), B₆₆ (-656); rms deviation between 51 experimental and calculated levels is 11 cm⁻¹.

^c Three largest terms of the percent free-ion mixture.

^d Theoretical free-ion centroids of ²⁵⁺¹L_j manifolds using set of atomic parameters listed in section 6.

^e Symmetry labels assigned with assistance from the calculation; data in tables 3 to 5 are compatible with these assignments.

^f Original levels used in calculation in parentheses: 2618 (2613), 5141 (5121), 6831 (6812).

5. Fluorescence and Selective Excitation Spectra

Fluorescence and selective excitation experiments at 1.6 K further assist in the assignment of Stark levels. The results of emission from the 1D_2 and 3P_0 levels are given in table 3.

The $16,409\text{ cm}^{-1}$ Stark level of the 1D_2 manifold was excited directly using a single-frequency cw dye laser of 2-MHz bandwidth. The result was the elimination of a number of weak fluorescence peaks that are observed by nonselective excitation. This is important since the $16,381\text{ cm}^{-1}$ transition which terminates on the Z_2 (19 cm^{-1}) level has comparable intensity to peaks associated with Pr^{3+} ions in other sites. Notwithstanding the narrow excitation bandwidth, several weak peaks were still observed, showing that perturbed sites with small energy shifts are present.

Figure 9, the emission from 1D_2 to the 3H_4 manifold at 1.6 K, shows three sharp peaks and two bands, each of which contains a reasonably sharp peak. The three sharp peaks are assigned as transitions to the levels Z_1 (0), Z_2 (19), and Z_3 (50), all in cm^{-1} . If the symmetry of each of these levels is Γ_3 , Γ_1 , and Γ_4 , respectively, then the originating Stark level in 1D_2 has Γ_2 symmetry. This assignment within the 1D_2 manifold is consistent with analysis of the hot bands observed in absorption. The relatively sharp peak in each of the two bands is used to establish additional Stark levels at 533 cm^{-1} (Γ_1) and 742 cm^{-1} (Γ_1). The broader structure is assigned to phonon-assisted transitions.

The 3P_0 level was excited by the 476.5-nm Ar-ion laser line and subsequent nonradiative relaxation. Consequently, several sites were involved in the emission. However, by comparing the derived energy levels with those obtained from the 1D_2 emission and with energy levels reported earlier by Hooge [7], we could assign the emission spectrum to levels of Pr^{3+} ions in D_2 sites for the 3H_4 and 3H_5 manifolds (see table 3). Levels derived from the 1D_2 emission not observed in the 3P_0 emission were assigned to Γ_1 symmetry based on electric-dipole selection rules for D_2 symmetry.

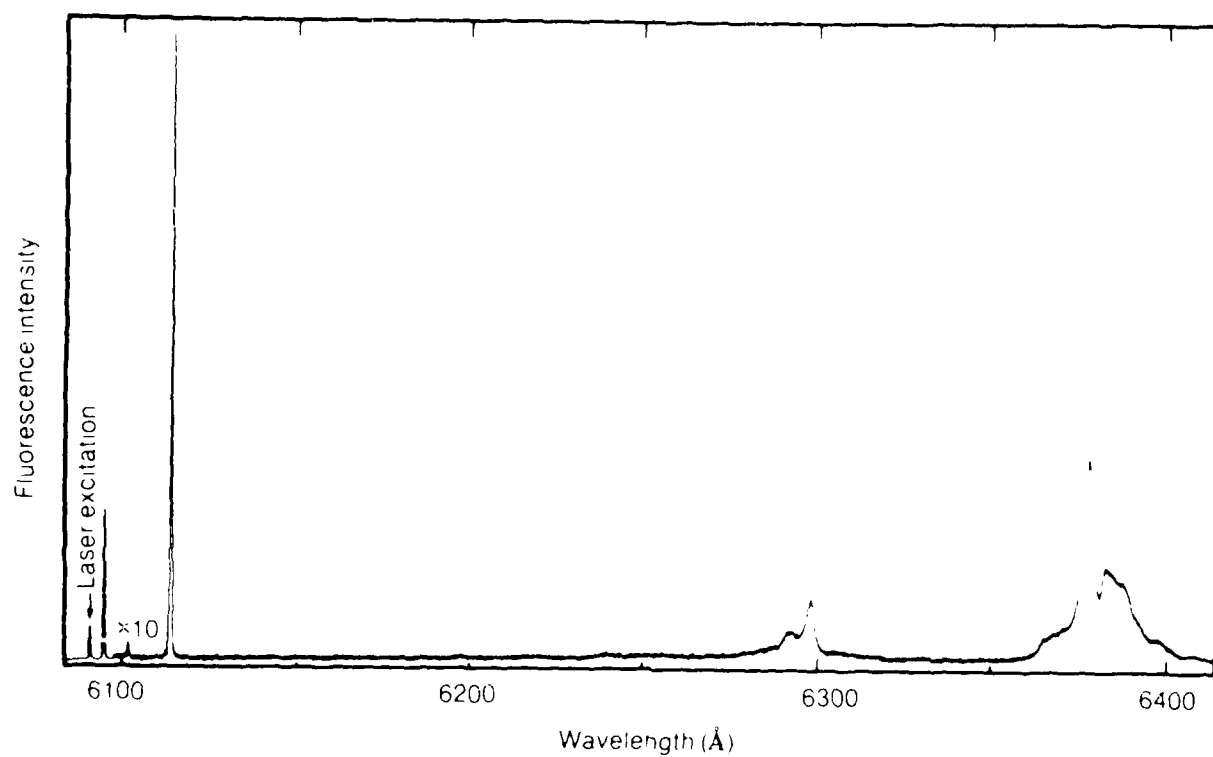


Figure 9. Site-selective laser excitation in 1D_2 manifold recorded at 1.6 K; emission is observed to 3H_4 ground state manifold.

6. Calculation of Crystal-Field Splitting

From the preceding analyses we have determined assignments for 17 Γ_1 , 12 Γ_2 , 10 Γ_3 , and 12 Γ_4 experimental Stark levels. These 51 levels were compared with the results obtained from the following crystal-field splitting calculation.

The free-ion wave functions were determined by diagonalizing a Hamiltonian containing the Coulomb interaction described in terms of Racah parameters $E^{(1)}$, $E^{(2)}$, and $E^{(3)}$ and the spin-orbit interaction parameter ζ . Also included in the free-ion calculation was the inter-configuration interaction in terms of parameters α , β , and γ . All parameters were expressed in units of cm^{-1} . The values of the free-ion parameters used in our calculation were obtained from Carnall et al [22] and are

$$E^{(1)} = 4548.2,$$

$$E^{(2)} = 21.935,$$

$$E^{(3)} = 466.73,$$

$$\zeta = 740.75,$$

$$\alpha = 21.255,$$

$$\beta = -799.94, \text{ and}$$

$$\gamma = 1342.9 \text{ (all in } \text{cm}^{-1}\text{)}.$$

The crystal-field splitting was calculated using the Hamiltonian [23]

$$H_{CF} = \sum_{i,k,m} B_{km} C_{km}(i), \quad (1)$$

where

$$C_{km}(i) = \sqrt{\frac{4\pi}{2k+1}} Y_{km}(\theta_i, \phi_i) \quad (2)$$

and the sum over i covers the two electrons of the configuration $4f^2$ of Pr^{3+} . For D_2 symmetry there are nine nonvanishing B_{km} crystal-field parameters with even k . For the yttrium site in YAG there are six nonequivalent magnetic sites which are related by rotation operations. Explicit algebraic expressions for each set of B_{km} parameters associated with one of the six equivalent sets are given by Morrison and Leavitt [4]. We use the first set as defined by them (page 633, ref. 4).

The first determination of a set of B_{km} parameters was obtained from a lattice-sum calculation that included point-charge, point-dipole, and self-induced contributions [24,25]. Table 6 gives the ion positions in the lattice, effective ionic electric charges, and the polarizability of the oxygen ions [26]. The parameters A_{km} (table 7) are related to the B_{km} parameters through the expression

$$B_{km} = \rho_k(Pr) A_{km} , \quad (3)$$

where $\rho_k(Pr)$ are radial factors [25]. The lattice-sum results provided an initial view of the overall splitting before the B_{km} parameters were varied to improve agreement with experimental energy levels and symmetry assignments.

Next, an empirical set of B_{km} parameters was determined from the best quadratic fit of B_{km} parameters obtained from earlier analyses of the observed spectra of Nd^{3+} , Tb^{3+} , Dy^{3+} , and Er^{3+} in YAG [24]. The calculated splitting using these parameters was in reasonable agreement with the splitting observed in $\text{Pr}^{3+}:\text{YAG}$. Subsequent variation of the parameters affected the symmetry assignment for only a few Stark levels. From these calculations, Z_3 (50 cm^{-1}) was assigned Γ_4 symmetry. This choice proved to be consistent in analyzing the 50-cm^{-1} hot bands and was the only exception to our algorithm of assigning symmetry labels to observed Stark levels.

Using established Γ_1 and Γ_3 experimental levels in the calculation, we allowed the B_{km} parameters to vary to obtain the best overall agreement between calculated and experimental levels. With the new set of parameters, we then predicted the splitting of the Γ_2 and Γ_4 levels and compared these results with our experimental levels. This calculation again predicted Z_3 (50 cm^{-1}) to have Γ_4 symmetry. A subsequent calculation including all 51 levels with their experimentally established symmetry labels was performed, giving a best overall rms deviation of 11 cm^{-1} . An additional six Stark levels, whose symmetry could not be established from experiment without some ambiguity, were identified through the results of the calculation. Table 7 compares the calculated and experimental levels for all manifolds.

7. Comments

With few exceptions, predicted splittings and symmetry assignments for the 3H_4 , 3H_5 , and 1D_2 manifolds are in agreement with levels and assignments established from emission and hot-band absorption data. Fluorescence from the 3P_0 manifold to a 576-cm^{-1} Stark level indicates that the level has Γ_2 , Γ_3 , or Γ_4 symmetry. Small changes in B_{km} parameters greatly affect the predicted splitting of 3H_4 . The ground-state and excited Stark levels at 19 cm^{-1} and 50 cm^{-1} are observed in both emission and absorption. The Γ_n labels deduced from the emission spectra agree with those deduced from the absorption spectra. Our results differ in some details from other results reported recently [9,11], in which higher concentrations of Pr^{3+} or higher temperatures were used.

Level Y_5 in the 3H_5 manifold is predicted to have Γ_2 symmetry. This is in agreement with an analysis of the hot bands observed in the absorption spectrum. However, the fluorescence spectra suggest a Γ_4 symmetry. Since 3H_5 has only three Γ_4 levels, we chose the three levels for which emission, absorption, and the crystal-field splitting calculation were all in agreement (namely Y_2 , Y_8 , and Y_{10}) as Γ_4 levels. We assigned Y_5 to Γ_2 symmetry in agreement with absorption data and the prediction by the crystal-field splitting calculation.

Emission from 3P_0 and 1D_2 to the 3H_6 manifold is generally very weak [7,9]. Hooge [7] reports fluorescence from 3P_0 to Stark levels at 4296 , 4333 , 4396 , 4451 , and 4561 cm^{-1} . These levels correspond to levels established from our absorption data as follows: 4305 (Γ_4), 4339 (Γ_3), 4409 (Γ_2), 4440 (Γ_3), and 4567 (Γ_4) cm^{-1} . The selection rules are consistent in analyzing both the emission and absorption data for 3H_6 . The crystal-field splitting calculation (table 7) is also in agreement with these assignments, although the 4339 (Γ_3) cm^{-1} level is predicted to lie somewhat higher. Levels not observed in emission from 3P_0 but identified from hot-band absorption data include 4316 (Γ_1), 4358 (Γ_1), 5045 (Γ_1), and 5141 (Γ_2) cm^{-1} . These levels agree with the calculated splitting given in table 7.

A small change in calculated splitting of the six highest energy Stark levels of the 3H_6 manifold has a great influence on the calculated splitting of the 3F_2 manifold. Unfortunately, we have experimental data on only three of these levels: 4843 (Γ_2), 5045 (Γ_1), and 5141 (Γ_2) cm^{-1} . Consequently, we reexamined our hot-band data and the reported levels for the 3F_2 manifold [7,9]. There are several intense hot

bands at 5379 and at 5535 cm^{-1} that persist even in the 15 K absorption spectrum. Furthermore, a pair of hot bands separated by 31 cm^{-1} implies a Γ_3 level at 5535 cm^{-1} , a level reported by others [7,9]. We chose W_1 (5365 cm^{-1}) to have Γ_4 symmetry since we found no evidence for a 50- cm^{-1} hot band. The calculation predicts a Γ_4 level at 5388 cm^{-1} . Hot-band data indicate that W_2 (5430 cm^{-1}) has Γ_2 symmetry and that W_5 (5585 cm^{-1}) has Γ_1 symmetry, in agreement with the calculated splitting.

The experimental assignments given to Stark levels of the 3F_3 and 3F_4 manifolds are in agreement with the calculated splitting. With only a few minor exceptions, our levels agree with those reported by Hooze [7]. The predicted splitting of the 1G_4 manifold was the most difficult to reconcile with experiment, since we could identify only some of the Stark levels out of the nine expected in D_2 symmetry. The predicted manifold splitting is over 1200 cm^{-1} (see table 7), with the four lowest energy levels found within 33 cm^{-1} of each other. A pair of hot bands separated by 31 cm^{-1} places a Γ_3 level at 10,380 cm^{-1} . The level observed at 9828 cm^{-1} appears to have Γ_1 symmetry, in agreement with Hooze [7]. The predicted level is nearly 80 cm^{-1} away from the observed level. We could not reduce this difference without affecting the overall agreement obtained for other established experimental levels used in the calculation.

The predicted splitting of the 1D_2 manifold is over 700 cm^{-1} . All five expected Stark levels and their symmetries have been identified from experiment; the calculated splitting is in reasonable agreement. Throughout all variations of B_{km} parameters, the two lowest Stark levels, having symmetry labels Γ_1 and Γ_2 , would interchange positions as the lowest energy level in the manifold. Site-selective excitation from the 16,409 cm^{-1} level is consistent with the assignment of Γ_2 symmetry, which is based on an analysis of the hot-band data in absorption.

Table 7 indicates that the predicted splitting of the 1I_6 manifold overlaps the 3P_1 manifold. Our assignments to the 3P_0 and 3P_1 manifolds agree with work reported earlier [7,9]. The limited number of observed spectra associated with the 1I_6 precluded an analysis of the splitting of this manifold. Assignments given to the Stark levels of the 3P_2 manifold are in agreement with results from the crystal-field splitting calculation.

8. Conclusions

From an analysis of the emission and absorption spectra of $\text{Pr}^{3+}:\text{YAG}$, a number of Stark levels have been assigned to Pr^{3+} ions occupying Y^{3+} ion sites of D_2 symmetry in the garnet lattice. The most intense spectra can be analyzed with consistency. However, the origin of the weak spectra is uncertain, since some Pr^{3+} ions are found in minority sites. On the basis of 51 Stark levels assigned to Pr^{3+} ions in D_2 sites, comparison between experimental and calculated levels using the B_{km} parameters given in table 7 yields an rms deviation of 11 cm^{-1} .

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